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INTERNATIONAL ASTRONOMICAL UNION**

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URL <http://cfa-www.harvard.edu/iau/cbat.html> ISSN 0081-0304  
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*POSSIBLE SUPERNOVA IN NGC 4945*

C. Jacques and E. Pimentel, Belo Horizonte, Brazil, report their discovery of an apparent supernova (mag 12.8) on unfiltered CCD images taken on Feb. 8.22 and 10.23 UT with a 0.30-m Schmidt-Cassegrain reflector in the course of the CEAMIG/REA Supernovae Search. The new object is located at  $\alpha = 13^{\text{h}}04^{\text{m}}44^{\text{s}}.06$ ,  $\delta = -49^{\circ}33'59''.8$  (equinox 2000.0), which is  $407''$  west and  $351''$  south of the center of NGC 4945, but which also lies in a rather rich area of foreground Milky Way stars in Centaurus. Nothing is visible on a CCD image taken by Jacques on 2004 June 20.93 (limiting mag 18.5) or on a red Digitized Sky Survey plate from 1976.

*SUPERNOVAE 2005Q, 2005S, 2005Y, 2005Z, 2005ad*

N. Morrell, M. Hamuy, and G. Folatelli, Carnegie Supernova Project; and F. Olivares, University of Chile, report that spectra (range 380–930 nm) were obtained on Feb. 4 and 8 UT with the Las Campanas 2.5-m du Pont Telescope (+ WFCCD spectrograph), showing that SNe 2005Q, 2005Y, 2005Z, and 2005ad (cf. *IAUC* 8473, 8476, 8479) are all of type II. SNe 2005ad, 2005Z, and 2005Q display blue continua with strong P-Cyg profiles in the Balmer lines and He I 587.6-nm, indicative of young events. SN 2005ad seems to be very young, based on the very blue continuum. SN 2005Y shows the Balmer lines and the Fe II 516.9-nm absorption, implying a more evolved event. The expansion velocities, derived from the minimum of the H $\beta$  absorption (assuming for the parent galaxies the recession velocities listed in the NED database) are: 2005Q, 11500 km/s; 2005Y, 6800; 2005Z, 11600; 2005ad, 7400. A spectrum of SN 2005S reveals it to be a type-Ia supernova, 2–3 weeks after maximum light.

*SUPERNOVAE 2005R, 2005X, AND 2005Y*

R. J. Foley and A. V. Filippenko, University of California, Berkeley, report that inspection of CCD spectra (range 500–1000 nm), obtained on Feb. 11 UT with the Keck II 10-m telescope (+ DEIMOS), shows that SN 2005R (*IAUC* 8473) is of type II $n$ ; the broad H $\alpha$  emission line has a full-width-at-half-maximum of  $\sim 2400$  km/s. SN 2005X (*IAUC* 8476) is of type Ia,  $\sim 2$  weeks past maximum brightness; the redshift of its host galaxy is measured to be  $z = 0.0745$ . SN 2005Y (*IAUC* 8476) is of type II, showing a prominent H $\alpha$  P-Cyg profile. After removing the host-galaxy recession velocity of 4900 km/s (from NED), the minimum of the broad H $\alpha$  absorption is measured to be  $-7400$  km/s relative to its rest wavelength.