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URL http://cfa-www.harvard.edu/iau/cbat.html ISSN 0081-0304 Phone 617-495-7440/7244/7444 (for emergency use only)

## COMETS C/2006 A4-A7, B2-B4, E2, F3 (SOHO)

Additional comets have been found on SOHO website images (cf. *IAUC* 8679) — all being Kreutz sungrazers except for C/2006 B4 (Meyer group) and for C/2006 E2 and C/2006 F3 (Marsden group). C/2006 A4 is described by K. Battams as very faint and diffuse. C/2006 A5 had a bright coma that "evaporated" quickly in the C2 images, with a tail that "lingered for a long time", being longer than 1° on Jan. 5.612 UT at ~  $4R_{\odot}$ ; in C3 images, the comet was still brightening when at mag perhaps -1.5 on Jan. 5.112 at  $11.6R_{\odot}$ , showing a bright, thin tail, but it was at mag ~ 2.9 and fading fast by the time it reached  $7R_{\odot}$ . The other comets were stellar in appearance. Peak magnitudes: C/2006 A6, 6.3, Jan. 10.431,  $10.8R_{\odot}$ ; C/2006 B2 and C/2006 F3 were very small and faint. C/2006 B4, at mag  $\approx$  7, was slightly elongated in the direction of motion. C/2006 E2 was extremely faint and slightly diffuse in C2 images (stellar in C3).

Comet	2006 UT	$\alpha_{2000}$	$\delta_{2000}$	Inst.	F	MPEC
C/2006 A4	Jan. 2.9	$19^{h}04$ $19^{h}04$	-25°26'	C3/2	$\mathbf{R}\mathbf{M}$	2006-F43
C/2006 A5	4.0	96 19 15.1	-2753	C3/2	TH	2006 - F43
C/2006 A6	10.0	19 40.1	$-24 \ 21$	C3	TH	2006-F43
C/2006 A7	12.4	31 19 49.2	$-24\ 12$	C3	$\mathbf{SF}$	2006 - F43
C/2006 B2	20.3	$21  20 \ 22.8$	-21 51	C3	HS	2006-F44
C/2006 B3	21.6	54 20 38.1	$-23\ 16$	C3	$\mathbf{R}\mathbf{K}$	2006-F44
C/2006 B4	26.4	96 20 37.2	$-17\ 18$	C2	HS	2006-F44
C/2006 E2	Mar. 15.0	$96  23 \ 50.5$	-126	C3/2	RM	2006 - F50
C/2006 F3	26.1	01 0 23.1	+ 4 13	C2	TH	2006 - F50

## V1647 ORIONIS

V. Venkat and B. G. Anandarao, Astronomy and Astrophysics Division, Physical Research Laboratory, Ahmedabad, report near-infrared photometry of V1647 Ori undertaken at Mt. Abu Observatory's 1.2-m telescope (+ NICMOS infrared array) on Feb. 27, revealing magnitudes J = 14.6, H = 11.7, and K' = 9.7. These magnitudes are close to the 2MASS values and fainter than their own values from 2005 Dec. 3 (Ojha *et al.* 2006, *MNRAS*, in press; http://xxx.lanl.gov/abs/astro-ph/?0602044). It appears therefore that, following its EXor-like eruption (*IAUC* 8289) that lasted more than 24 months (cf. *IAUC* 8600; 8681), the star may be near its pre-eruptive state.

2006 March 29

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