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## COMET P/2006 W4 = P/1993 D1 (HILL)

S. Foglia, R. Matson, and M. Tombelli report the identification of "precovery" images of comet P/2006 W4 (cf. *IAUC* 8779), aided by the orbital elements published on *MPEC* 2007-X14 (with indicated  $\Delta T = +1.06$  days when run back to 1992), and their astrometric measurements are provided below from the trails on the two U.K. Schmidt Telescope plates.

1993 UT	$lpha_{2000}$	$\delta_{2000}$
Feb. 26.53282	$9^{h}44^{m}59.08$	$-34^{\circ}38^{'}11\overset{''}{.0}$
26.57795	$9\ 44\ 57.60$	$-34 \ 38 \ 01.8$
Apr. 5.58208	$13 \ 04 \ 38.14$	$-37\ 52\ 27.4$
-5.62375	$13 \ 04 \ 36.85$	-37 52 15.9

The following orbital elements by B. G. Marsden are from 92 observations, 1993–2007 (mean residual 0".6):

Epoch = $1992$ June 27.0 TT			
	T = 1992 June 21.6666 TT	$\omega = 247.5818$	
	e = 0.314858	$\Omega = 243.4600 $ 2000.0	
	q = 4.416239  AU	i = 36.2938 J	
a	$n = 6.445725 \text{ AU} \qquad n^{\circ} = 0.0602$	P = 16.36  years	
Epoch = $2009$ Jan. $9.0$ TT			
	T = 2009 Jan. 21.0916 TT	$\omega = 249^{\circ}.6184$	
	e = 0.314866	$\Omega = 243.2496 $ 2000.0	
	q = 4.438583  AU	i = 36.3618 J	
a	$n = 6.478413 \text{ AU}  n^{\circ} = 0.0597$	P = 16.49  years	

## V597 PUPPIS

R. J. Rudy, R. W. Russell, and D. K. Lynch, The Aerospace Corporation; and C. E. Woodward, University of Minnesota, report 0.8- to 2.5- $\mu$ m spectroscopy of this nova (cf. *IAUC* 8895, 8896) using the Infrared Telescope Facility (+ SpeX) on Nov. 30.54 UT. V597 Pup was in its early stage of development. He I 1.0830- $\mu$ m was very strong, as was He I 20581- $\mu$ m. The O I lines were strong, and Br $\gamma$  showed a doubled line profile. There was no evidence of He II lines or thermal emission from dust. The FWHM of the lines was  $\approx$  3900 km/s.

Visual magnitude estimates by A. Amorim, Florianopolis, Brazil: Nov. 16.047 UT, 8.1; 20.041, 9.7; 22.078, 9.8; 23.040, 10.0.

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